# EXAMINATION TENSOR CALCULUS AND DIFFERENTIAL GEOMETRY

Course code: 2WAH0. Date: Tuesday June 26, 2018. Time: 18h00-21h00. Place: AUD 12

#### Read this first!

- Write your name and student identification number on each paper.
- The exam consists of 3 problems. The maximum credit for each item is indicated in the margin.
- You may consult an immaculate hardcopy of the online draft notes "Tensor Calculus and Differential Geometry (2WAH0)" by Luc Florack. No other material or equipment may be used.



## (20) 1. TENSOR CALCULUS MISCELLANY.

Consider the mixed tensor  $\mathbf{X} = X_j^i \mathbf{e}_i \otimes \hat{\mathbf{e}}^j$ , and a basis transformation  $\mathbf{e}_i = A_i^j \mathbf{f}_j$ , together with its induced dual basis transformation,  $\hat{\mathbf{e}}^i = B_j^i \hat{\mathbf{f}}^j$ . The transformed holor  $\overline{X}_j^i$  is defined via  $\mathbf{X} = \overline{X}_j^i \mathbf{f}_i \otimes \hat{\mathbf{f}}^j$ .

(5) **a1.** Show that  $A_i^k B_k^j = \delta_i^i$ .

We have  $\delta_i^j = \langle \hat{\mathbf{e}}^j, \mathbf{e}_i \rangle = \langle B_k^j \hat{\mathbf{f}}^k, A_i^\ell \mathbf{f}_\ell \rangle = A_i^\ell B_k^j \langle \hat{\mathbf{f}}^k, \hat{\mathbf{f}}_\ell \rangle = A_i^\ell B_k^j \delta_\ell^k = A_i^k B_k^j$ .

(5) **a2.** Show that if the holor is invariant, i.e. if  $\overline{X}_{j}^{i} = X_{j}^{i}$  for any basis transformation, then **X** must equal the Kronecker tensor up to a constant factor.

On the one hand we have the general 'tensor transformation law'  $\overline{X}_{j}^{i} = A_{k}^{i}B_{j}^{\ell}X_{\ell}^{k}$ . On the other hand we have the invariance requirement  $\overline{X}_{j}^{i} = X_{j}^{i}$ . Combining these, and using a, it follows that  $A_{k}^{j}X_{j}^{i} = A_{j}^{i}X_{k}^{j}$  regardless of the choice of the transformation matrix. We may therefore regard the  $A_{j}^{i}$  as independent variables, and differentiate the equation with respect to these. This yields  $X_{j}^{i}\delta_{q}^{p} = X_{q}^{p}\delta_{j}^{i}$ . Contraction over p = q produces  $nX_{j}^{i} = X\delta_{j}^{i}$ , in which  $X = X_{k}^{k}$ , in other words  $X_{j}^{i} = c\delta_{j}^{i}$ , in which c = X/n is an arbitrary constant. We conclude that  $\mathbf{X} = c\delta_{i}^{i} \mathbf{e}_{i} \otimes \hat{\mathbf{e}}^{j} = c \mathbf{e}_{i} \otimes \hat{\mathbf{e}}^{i} = c \langle -, - \rangle$  is a multiple of the Kronecker tensor.

Next, consider the tensors  $\mathbf{V} = V^i \mathbf{e}_i$ ,  $\mathbf{S} = S^{ij} \mathbf{e}_i \otimes \mathbf{e}_j$  and  $\mathbf{T} = T_{ijk} \, \hat{\mathbf{e}}^i \otimes \hat{\mathbf{e}}^j \otimes \hat{\mathbf{e}}^k$ .

(5) **b1.** Expand the holor  $V^{(i}S^{jk)}$  of  $\mathscr{S}(\mathbf{V} \otimes \mathbf{S})$ , i.e. the symmetrised outer product of  $\mathbf{V}$  and  $\mathbf{S}$ , explicitly in terms of the holors  $V^i$  and  $S^{jk}$ .

We have

$$V^{(i}S^{jk)} \doteq \mathscr{S}(\mathbf{V} \otimes \mathbf{S})^{ijk} = \frac{1}{6} \left( V^{i}S^{jk} + V^{i}S^{kj} + V^{j}S^{ik} + V^{j}S^{ki} + V^{k}S^{ij} + V^{k}S^{ji} \right) \,.$$

(5) **b2.** Suppose  $\mathscr{S}(\mathbf{V} \otimes \mathbf{S}) = 0$  for all vectors  $\mathbf{V}$ . Show that  $\mathscr{S}(\mathbf{S}) = 0$ .

Differentiating the null function

$$0 = f^{ijk}(\mathbf{V}) \doteq V^i S^{jk} + V^i S^{kj} + V^j S^{ik} + V^j S^{ki} + V^k S^{ij} + V^k S^{ji} = 2\left(V^i S^{(jk)} + V^j S^{(ik)} + V^k S^{(ij)}\right),$$

with respect to  $V^{\ell}$  for fixed S yields

$$0 = \frac{1}{2} \frac{\partial f^{ijk}(\mathbf{V})}{\partial V^{\ell}} = \delta^i_{\ell} S^{(jk)} + \delta^j_{\ell} S^{(ik)} + \delta^k_{\ell} S^{(ij)} \,.$$

Contraction of the indices i and  $\ell$  yields  $(n+2)S^{(jk)} = 0$ , in which n is the dimension of the vector space, whence  $S^{(jk)} = 0$ .

#### (30) 2. LAPLACE-DE RHAM OPERATOR.

The Laplace-de Rham operator  $\Delta_{LDR}$ , a generalization of the Laplace-Beltrami operator  $\Delta_{LB}$ , is defined so as to act on sections of any k-form bundle over a Riemannian manifold M of dimension  $n, 0 \le k \le n$ . Its formal definition is

$$\Delta_{\rm LDR} \doteq d\delta + \delta d \,,$$

in which d denotes the exterior derivative or differential, and  $\delta$  the so-called *codifferential*. The latter is defined for a k-form field in terms of the differential d and Hodge star operator \* as follows:

$$\delta = (-1)^{(k+1)n+1} * d * .$$

For notational simplicity we denote the linear space of k-form fields by  $\Lambda_k(\mathbf{M})$ ,  $0 \le k \le n$ . Recall that for a k-form field  $\boldsymbol{\omega} = \omega_{i_1...i_k} dx^{i_1} \land \ldots \land dx^{i_k} \in \Lambda_k(\mathbf{M})$ , say, the differential  $d\boldsymbol{\omega} \in \Lambda_{k+1}(\mathbf{M})$  is defined as  $d\boldsymbol{\omega} = \partial_i \omega_{i_1...i_k} dx^{i_1} \land \ldots \land dx^{i_k}$ , thus d increments covariant rank by one.

Terminology. A 0-form field is usually referred to as a *scalar field*, an *n*-form field as a *volume form*.

In this problem you may use the following lemma without proof.

**Lemma.** The double Hodge star \*\* preserves the covariant rank of its operand. More specifically, on a (positive definite) Riemannian manifold M of dimension n we have:

$$** = (-1)^{k(n-k)} \mathbf{1}_{\Lambda_k(\mathbf{M})},$$

in which  $1_{\Lambda_k(\mathbf{M})}$  is the identity operator (usually suppressed in the notation by writing  $** = (-1)^{k(n-k)}$ ).

- (5) **a1.** Show that for a scalar field  $f \in \Lambda_0(\mathbf{M})$  we have  $\Delta_{\text{LDR}} f = \delta df$ .
- (5) **a2.** Show that for a volume form  $\rho \in \Lambda_n(M)$  we have  $\Delta_{LDR}\rho = d\delta\rho$ .

By definition  $*1 = \epsilon \in \Lambda_n(M)$ , the unit *n*-form. Since the differential *d* increments covariant order of its operand, we have  $d\epsilon \in \Lambda_{n+1}(M)$ , and therefore  $d\epsilon = 0$ . By linearity we then also have  $d * f = fd * 1 = fd\epsilon = 0$ , leading to the reduction  $\Delta_{\text{LDR}} f = (d\delta + \delta d)f = \delta df$ . By the same token,  $d\rho \in \Lambda_{n+1}(M)$  and thus  $d\rho = 0$ , whence  $\Delta_{\text{LDR}} \rho = (d\delta + \delta d)\rho = d\delta\rho$ .

(5) **b.** Argue that if  $\omega \in \Lambda_k(M)$ , then  $\delta \omega \in \Lambda_{k-1}(M)$ , i.e.  $\delta$  decrements covariant rank by one. [*Hint:* Avoid explicit calculation of  $\delta \omega$ .]

Use the prototypes of  $*: \Lambda_k(\mathbf{M}) \to \Lambda_{n-k}(\mathbf{M})$  and  $d: \Lambda_k(\mathbf{M}) \to \Lambda_{k+1}(\mathbf{M})$  for any  $0 \le k \le n$ . Ignoring the irrelevant sign, we thus have  $\delta(k$ -form) = (\*d\*)(k-form) = (\*d)((n-k)-form) = \*((n-k+1)-form) = (n-(n-k+1))-form = (k-1)-form.

(5) **c.** Show that we may also write the definition of the Laplace-de Rham operator as  $\Delta_{LDR} = (d + \delta)^2$ .

This boils down to proving that  $d^2 = \delta^2 = 0$ .

Proof of  $d^2 = 0$ : Let  $\boldsymbol{\omega} \in \Lambda_k(\mathbf{M})$  be a section of the k-form bundle, say  $\boldsymbol{\omega} = \omega_{i_1...i_k} dx^{i_1} \wedge \ldots \wedge dx^{i_k}$ . Then  $d\boldsymbol{\omega} \in \Lambda_{k+1}(\mathbf{M})$  is defined as  $d\boldsymbol{\omega} = \partial_i \omega_{i_1...i_k} dx^i \wedge dx^{i_1} \wedge \ldots \wedge dx^{i_k}$ , and  $d^2 \boldsymbol{\omega} = \partial_j \partial_i \omega_{i_1...i_k} dx^j \wedge dx^i \wedge dx^{i_1} \wedge \ldots \wedge dx^{i_k} = 0$ . The latter follows from symmetry considerations w.r.t. the (i, j)-sum:  $\partial_j \partial_i$  is symmetric, whereas  $dx^j \wedge dx^i$  antisymmetric.

Proof of  $\delta^2 = 0$ :  $\delta^2 \propto (*d*)(*d*) \stackrel{\dagger}{\propto} *d^2* \stackrel{\ddagger}{=} 0$ , in which  $\propto$  indicates equality up to a constant (sign) factor. In the step marked by  $\dagger$  we have used the fact that  $** \propto 1$  is the identity operator up to a possible minus sign\*. In the step marked by  $\ddagger$  we have used  $d^2 = 0$ .

\*As an aside, the exact form of the double Hodge star in a pseudo-Riemannian manifold with metric determinant  $g \neq 0$  is given by  $** = (-1)^{k(n-k)} \operatorname{sgn} g$ . The codifferential then also acquires a possible minus sign:  $\delta = (-1)^{(k+1)n+1} \operatorname{sgn} g * d*$ .

For a scalar field f the Laplace-Beltrami operator  $\Delta_{\text{LB}}$  is defined as

$$\Delta_{ ext{lb}}f\doteqrac{1}{\sqrt{g}}\partial_i\left(\sqrt{g}g^{ij}\partial_jf
ight)\,,$$

with g the determinant of the Gram matrix with entries  $g_{ij}$ , and  $g^{ij}$  the holor of the inverse Gram matrix.

(10) **d.** Derive the relation between  $\Delta_{\text{LDR}} f$  and  $\Delta_{\text{LB}} f$  for a scalar field f. [*Hint:* (i)  $*(\omega_k dx^k) = g^{k\ell} \omega_k \epsilon_{\ell i_2...i_n} dx^{i_2} \otimes ... \otimes dx^{i_n}$ ; (ii) \*\*1 = 1; (iii)  $\epsilon_{i_1...i_n} = \sqrt{g} [i_1, ..., i_n]$ .]

We have  $\Delta_{\text{LDR}} f = \delta df = -*d*df$ . Notice the minus sign, which originates from the definition of  $\delta$  acting on a 1-form (viz. df). Working this out from right to left we first obtain

$$df = \partial_k f dx^k$$

Taking the Hodge star yields, using the hint,

$$*df = (g^{k\ell}\sqrt{g}\partial_k f)\mu_{\ell i_2...i_n}dx^{i_2}\otimes...\otimes dx^{i_n}$$

in which the notation  $\mu_{i_1...i_n} = [i_1, ..., i_n]$  has been used merely to maintain manifest covariant indices. The factor  $\sqrt{g}$  has been factored out so as to clarify the *x*-dependence of the holor of this (n-1)-form, viz. the part in parentheses. Taking the differential yields

$$d * df = \partial_m \left( g^{k\ell} \sqrt{g} \partial_k f \right) \mu_{\ell i_2 \dots i_n} dx^m \otimes dx^{i_2} \otimes \dots \otimes dx^{i_n}$$

Putting back the original holor of the  $\epsilon$ -tensor,  $\sqrt{g} \mu_{\ell i_2...i_n} = \epsilon_{\ell i_2...i_n}$ , we get

$$d * df = \frac{1}{\sqrt{g}} \partial_m \left( g^{k\ell} \sqrt{g} \partial_k f \right) \epsilon_{\ell i_2 \dots i_n} dx^m \otimes dx^{i_2} \otimes \dots \otimes dx^{i_n} .$$

Consider the tensor  $\epsilon_{\ell i_2...i_n} dx^m \otimes dx^{i_2} \otimes ... \otimes dx^{i_n}$ . Since all indices  $i_2, ..., i_n \in \{1, ..., n\}$  are effectively distinct, the only nontrivial option for the free index *m* results from the choice  $m = \ell$ , distinct from each  $i_2, ..., i_n$ . As a result we have

$$\epsilon_{\ell i_2 \dots i_n} dx^m \otimes dx^{i_2} \otimes \dots \otimes dx^{i_n} = \delta_\ell^m \epsilon_{i_1 \dots i_n} dx^{i_1} \otimes \dots \otimes dx^{i_n} = \delta_\ell^m \epsilon.$$

Consequently,

$$d * df = \frac{1}{\sqrt{g}} \partial_\ell \left( g^{k\ell} \sqrt{g} \partial_k f \right) \boldsymbol{\epsilon} \,.$$

Next, taking the Hodge dual once again, using  $*\epsilon = **1 = 1$ , we obtain

$$*d*df = \frac{1}{\sqrt{g}}\partial_\ell \left(g^{k\ell}\sqrt{g}\partial_k f\right) \,.$$

Finally, looking at the original definition, including a minus sign, we observe that

$$\Delta_{\rm LDR} f \doteq - * d * df \doteq -\Delta_{\rm LB} f \, .$$

## (50) 3. RIEMANN, RICCI & COTTON TENSOR.

Recall the definition of the (holor of the) *Riemann tensor* on an *n*-dimensional Riemannian manifold in terms of the Levi-Civita Christoffel symbols  $\Gamma_{ii}^k$ :

$$R_{i\ell j}^{k} = \partial_{\ell} \Gamma_{ij}^{k} - \partial_{j} \Gamma_{i\ell}^{k} + \Gamma_{\lambda\ell}^{k} \Gamma_{ij}^{\lambda} - \Gamma_{\lambda j}^{k} \Gamma_{i\ell}^{\lambda}.$$

The (holor of the) Ricci tensor is defined as

$$R_{ij} = R_{ikj}^k \,. \tag{(*)}$$

By definition, the fully covariant Riemann tensor has holor

$$R_{ijk\ell} = g_{im} R^m_{jk\ell} \,.$$

(5) **a.** Show that  $R_{ij} = R_{ji}$ . [*Hint:* Write  $R_{ij} = R_{ikj}^k = g^{k\ell} R_{\ell i k j}$  and observe that  $R_{\ell i k j} = R_{k j \ell i}$ .]

Following the hint we write

$$R_{ij} = R_{ikj}^{k} = g^{k\ell} R_{\ell ikj} = g^{k\ell} R_{kj\ell i} = R_{j\ell i}^{\ell} = R_{ji}.$$

The *Ricci scalar* is the trace of the Ricci tensor:  $R = g^{ij}R_{ij}$ . Henceforth,  $D_k$  refers to a covariant derivative corresponding to the Levi-Civita connection.

(5) **b.** Expand  $D_k R_{ij}$  in terms of (0<sup>th</sup> and 1<sup>st</sup> order) partial derivatives of  $R_{ij}$  and Christoffel symbols, but do *not* expand  $R_{ij}$  itself.

 $D_k R_{ij} = \partial_k R_{ij} - \Gamma_{ik}^{\ell} R_{\ell j} - \Gamma_{jk}^{\ell} R_{i\ell}.$ 

(5) **c.** Show that, in dimension n=1,  $R_{ijk\ell} = 0$ . [*Hint:* The covariant holor  $R_{ijk\ell}$  inherits a certain antisymmetry property from the mixed holor  $R_{ijk\ell}^m$ .]

By antisymmetry  $R_{ijk\ell} = -R_{ij\ell k}$  we have, in n = 1,  $R_{1111} = -R_{1111}$  as the single degree of freedom, whence  $R_{ijk\ell} = 0$ .

Without proof we state that the fully covariant Riemann tensor in n=2 has the following form:

$$n = 2: \qquad R_{ijk\ell} = (g_{ik}g_{j\ell} - g_{i\ell}g_{jk})\psi,$$

for some scalar field  $\psi \in C^{\infty}(\mathbf{M})$ .

(5) **d.** Show that, in this 2-dimensional case,  $\psi = \frac{1}{2}R$ .

We have  $R_{ij} = g^{k\ell} R_{\ell i k j} \stackrel{\star}{=} g^{k\ell} (g_{\ell k} g_{ij} - g_{\ell j} g_{ik}) \psi = (ng_{ij} - g_{ij}) \psi \stackrel{n \equiv 2}{=} g_{ij} \psi$ , whence  $R \stackrel{\text{def}}{=} g_{ij} R_{ij} = g^{ij} g_{ij} \psi = n \psi \stackrel{\star}{=} 2 \psi$ . The identities marked with  $\star$  exploit the conjecture for n = 2.

Likewise without proof we state that, in n = 3 dimensions,

$$n = 3: \qquad R_{ijk\ell} = a(g_{ik}R_{j\ell} - g_{jk}R_{i\ell} - g_{i\ell}R_{jk} + g_{j\ell}R_{ik}) + b(g_{ik}g_{j\ell} - g_{i\ell}g_{jk})R,$$

for certain constants  $a, b \in \mathbb{R}$ .

(5) **e.** Show that a = 1 and  $b = -\frac{1}{2}$ .

Contraction with  $g^{ik}$  yields  $R_{j\ell} \stackrel{\text{def}}{=} g^{ik} R_{ijk\ell} \stackrel{\star}{=} g^{ik} \left( a(g_{ik}R_{j\ell} - g_{jk}R_{i\ell} - g_{i\ell}R_{jk} + g_{j\ell}R_{ik}) + b(g_{ik}g_{j\ell} - g_{i\ell}g_{jk}) \right)$ . The identity marked with  $\star$  exploits the conjecture for n = 3. Simplifying the r.h.s. we obtain, using  $g^{ik}R_{ik} = R$  and  $g^{ik}g_{ik} = n = 3$  in this specific case:  $R_{j\ell} = (n-2)aR_{j\ell} + (a + (n-1)b)g_{j\ell}R \stackrel{n \equiv 3}{=} aR_{j\ell} + (a + 2b)g_{j\ell}R$ , whence a = 1 and  $b = -\frac{1}{2}$ .

We consider a conformal metric transformation,  $\tilde{g}_{ij} = e^{2\phi}g_{ij}$ , in which  $\phi \in C^{\infty}(\mathbf{M})$  is a smooth scalar field. The Levi-Civita Christoffel symbols associated with the transformed metric are indicated by  $\tilde{\Gamma}_{ij}^k$ .

(5) **f1.** Show that  $\widetilde{\Gamma}_{ij}^k = \Gamma_{ij}^k + S_{ij}^k$  and derive the explicit form of the symbols  $S_{ij}^k$  in terms of the scalar field  $\phi$  and the original Riemannian metric tensor  $g_{ij}$ .

We have  $S_{ij}^k = \partial_i \phi \delta_i^k + \partial_j \phi \delta_i^k - \partial^k \phi g_{ij}$ , in which  $\partial^k \phi = g^{k\ell} \partial_\ell \phi$ .

(5) **f2.** Argue why  $S_{ij}^k$ , as opposed to  $\Gamma_{ij}^k$ , is the holor of a tensor. [*Hint:* Avoid elaborate computations. Use the observations  $D_k g_{ij} = 0$  and  $D_k f = \partial_k f$  for any scalar field f.]

In the expression  $S_{ij}^k = \partial_i \phi \delta_j^k + \partial_j \phi \delta_i^k - \partial^k \phi g_{ij}$  we may, following the hint, replace each  $\partial$  by D, showing the tensorial nature of  $S_{ij}^k$ .

(5) **g1.** Show that in general dimension *n* the holor  $\widetilde{R}_{i\ell j}^k$  of the Riemann tensor after conformal metric transformation is given in terms of  $R_{i\ell j}^k$  and the tensorial holor  $S_{ij}^k$  as follows:

$$\widetilde{R}^k_{i\ell j} = R^k_{i\ell j} + D_\ell S^k_{ij} - D_j S^k_{i\ell} + S^k_{\lambda\ell} S^\lambda_{ij} - S^k_{\lambda j} S^\lambda_{i\ell}$$

Insert  $\widetilde{\Gamma}_{ij}^k = \Gamma_{ij}^k + S_{ij}^k$  into definition of  $\widetilde{R}_{i\ell j}^k$  in terms of  $\widetilde{\Gamma}$ -symbols, and the result follows.

(5) **g2.** Now take n=2. Find  $\widetilde{R}_{ij}$  in terms of  $R_{ij}$  and  $\phi$ .

Using the definition of the Ricci tensor and its simplification in n=2,  $R_{ij} = g^{k\ell}R_{\ell ikj} \stackrel{d}{=} \frac{1}{2}Rg_{ij}$ , with  $\Gamma_{ij}^k$  replaced by  $\widetilde{\Gamma}_{ij}^k = \Gamma_{ij}^k + S_{ij}^k$ , we obtain, using f1,  $\widetilde{R}_{ij} = R_{ij} - g_{ij}\Delta_{\text{LB}}\phi$ , in which  $\Delta_{\text{LB}}\phi = D^k D_k\phi = D^k \partial_k\phi$ .

Note: In general dimension n we have  $\widetilde{R}_{ij} = R_{ij} - g_{ij}\Delta_{LB}\phi - (n-2)(\partial_i\phi\partial_j\phi - D_i\partial_j\phi - \|\nabla\phi\|^2 g_{ij})$ . This general result follows from f1 and g1 after a more tedious but straightforward computation.

The so-called *Cotton tensor* is a third order covariant tensor with the following holor:

$$C_{ijk} = D_k R_{ij} - D_j R_{ik} + \frac{1}{2(n-1)} \left( g_{ik} D_j R - g_{ij} D_k R \right) \,.$$

(5) **h.** Show that in n=2 the Cotton vanishes identically:  $C_{ijk} = 0$ .

Using the n = 2 form of the fully covariant Riemann tensor we find  $C_{ijk} = \frac{1}{2}(g_{jk}D_jR - g_{ij}D_kR)$ . With the help of e we obtain  $R_{ij} = \frac{1}{2}Rg_{ij}$ . Inserting this in the definition of the Cotton tensor yields (remember to take n=2)  $C_{ijk} = 0$ .

THE END